

PROBLEM 1

Giada Selva S4763440

a) $ds^2 = -\left(1 - \frac{2M}{r}\right) dt^2 + \left(1 - \frac{2M}{r}\right)^{-1} dr^2 + r^2 (d\theta^2 + \sin^2\theta d\phi^2)$

is the Schwarzschild metric in geometrical units.

The metric is independent of t , the Killing vector is $\xi^\alpha = (1, 0, 0, 0)$.

there is also a manifest rotational symmetry, as the metric is independent of ϕ ,

with Killing vector $\eta^\alpha = (0, 0, 0, 1)$ and

$$\frac{\partial g_{\alpha\beta}}{\partial \phi} = 0 \quad \text{and} \quad \frac{\partial \mathcal{L}}{\partial \phi} = 0$$

Euler-Lagrange eq:

$$-\frac{d}{d\sigma} \left(\frac{\partial \mathcal{L}}{\partial \left(\frac{d\phi}{d\sigma} \right)} \right) + \frac{\partial \mathcal{L}}{\partial \phi} = 0 \rightarrow -\frac{d}{d\sigma} \left(\frac{\partial \mathcal{L}}{\partial \left(\frac{d\phi}{d\sigma} \right)} \right) = 0$$

and

$$\frac{\partial \mathcal{L}}{\partial \left(\frac{d\phi}{d\sigma} \right)} = \text{const}$$

$$\mathcal{L} = \left[-g_{\alpha\beta} \frac{dx^\alpha}{d\sigma} \frac{dx^\beta}{d\sigma} \right]^{\frac{1}{2}}$$

$$\frac{\partial \mathcal{L}}{\partial \left(\frac{d\phi}{d\sigma} \right)} = \frac{1}{2} (\mathcal{L}^2)^{-\frac{1}{2}} \left(-g_{3\beta} \frac{dx^\beta}{d\sigma} \right) = -\frac{1}{2\mathcal{L}} g_{3\beta} \frac{dx^\beta}{d\tau} \frac{\dot{\mathcal{L}}}{d\sigma} = \text{const}$$

$$\rightarrow g_{\alpha\beta} \eta^\alpha u^\beta = \underset{\wedge}{\eta} \cdot \underset{\wedge}{u} = \text{const} \rightarrow \underset{\wedge}{\eta} \cdot \underset{\wedge}{u} \text{ is conserved}$$

$$\eta \cdot u = g_{\phi\phi} \frac{d\phi}{d\tau} = r^2 \sin^2\theta \frac{d\phi}{d\tau} = r^2 \frac{d\phi}{d\tau} \quad \text{assuming } \theta(\tau) = \frac{\pi}{2}$$

the quantity $\ell = r^2 \frac{d\phi}{d\tau}$ is conserved along the geodesic.

it is the angular momentum per unit mass.

b) $\frac{M}{r^2} \left(\frac{dt}{d\tau} \right)^2 - r \left(\frac{d\phi}{d\tau} \right)^2 = 0$ is the equation of the geodesic

circular orbit of radius $R = 4M$

at $r = R$: $M \left(\frac{dt}{d\tau} \right)^2 - R^3 \left(\frac{d\phi}{d\tau} \right)^2$ $R = 4M$

$$\left(\frac{dt}{d\tau} \right)^2 - 64 M^2 \left(\frac{d\phi}{d\tau} \right)^2 = 0 \quad \rightarrow \quad \frac{dt}{d\tau} = 8M \frac{d\phi}{d\tau}$$

$$\Delta t = 8M \oint d\phi = 8M \int_0^{2\pi} d\phi = 16\pi M$$

$$\Delta t = 16\pi M$$

c) $u \cdot u = -1$ and $\theta = \frac{\pi}{2}$

$$g_{\alpha\beta} u^\alpha u^\beta = -1$$

$$g_{tt} \left(\frac{dt}{d\tau} \right)^2 + g_{rr} \left(\frac{dr}{d\tau} \right)^2 + g_{\phi\phi} \left(\frac{d\phi}{d\tau} \right)^2 = -1$$

$$-\left(1 - \frac{2M}{r}\right) \left(\frac{dt}{d\tau} \right)^2 + \left(1 - \frac{2M}{r}\right)^{-1} \left(\frac{dr}{d\tau} \right)^2 + r^2 \left(\frac{d\phi}{d\tau} \right)^2 = -1$$

on the geodesic $r = 4M = R$, $dr = 0$

$$-\left(1 - \frac{2M}{4M}\right) \left(\frac{dt}{d\tau} \right)^2 + 16M^2 \left(\frac{d\phi}{d\tau} \right)^2 = -1$$

from point b) : $\left(\frac{dt}{d\tau} \right)^2 = 64 M^2 \left(\frac{d\phi}{d\tau} \right)^2$

$$-\frac{1}{2} \cdot 64 M^2 \left(\frac{d\phi}{d\tau} \right)^2 + 16 M^2 \left(\frac{d\phi}{d\tau} \right)^2 = -1$$

$$16 M^2 \left(\frac{d\phi}{d\tau} \right)^2 = 1 \quad \rightarrow \quad \frac{d\phi}{d\tau} = \frac{1}{4M}$$

$$\Delta\tau = \oint d\tau = \int_0^{2\pi} \frac{d\tau}{d\phi} d\phi = 4M \int_0^{2\pi} d\phi = 8\pi M$$

$$\Delta\tau = 8\pi M$$

PROBLEM 2

$$\dot{a}^2 - \frac{8\pi\rho}{3} a^2 = 1$$

is the Friedmann equation, where

$\rho(t) = \rho(t_0) \left(\frac{a(t_0)}{a(t)} \right)^3$ is the total energy density

$$\rho(t_0) = \frac{3}{8\pi} H_0^2 \Omega, \quad H_0 = \frac{\dot{a}(t_0)}{a(t_0)}$$

here $k = -1, \Omega < 1$ that is negative curvature (open universe)

a) $\rho(t) = \frac{3}{8\pi} H_0^2 \Omega \left(\frac{a_0}{a} \right)^3$ where $a_0 = a(t_0)$

substituting in the Friedmann equation:

$$\dot{a}^2 - \frac{8\pi}{3} \cdot \frac{3}{8\pi} H_0^2 \Omega a^2 \frac{a_0^3}{a^3} = 1$$

$$\dot{a}^2 - H_0^2 \Omega \frac{a_0^3}{a} = 1$$

let $t = t_0$ and multiply with $\frac{1}{a_0^2}$

$$\left(\frac{\dot{a}_0}{a_0} \right)^2 - H_0^2 \Omega \frac{a_0^3}{a_0^3} = \frac{1}{a_0^2}$$

$$H_0^2 - H_0^2 \Omega = \frac{1}{a_0^2} \rightarrow a_0^2 H_0^2 (1 - \Omega) = 1 \quad \text{with } \Omega < 1$$

so we get:

$$a_0 = H_0^{-1} (1 - \Omega)^{-\frac{1}{2}}$$

b) let $A^2 = \frac{\Omega}{H_0 (1 - \Omega)^{3/2}}$

from point a):

$$\left. \begin{aligned} a_0 H_0 &= (1 - \Omega)^{-\frac{1}{2}} \\ \dot{a}^2 - a_0^3 \frac{H_0^3}{H_0} \frac{\Omega}{a} &= 1 \end{aligned} \right\} \rightarrow \dot{a}^2 - \frac{\Omega}{H_0 (1 - \Omega)^{3/2}} \cdot \frac{1}{a} = 1$$

and

$$\dot{a}^2 - \frac{A^2}{a} = 1 \rightarrow$$

$$\dot{a}^2 - 1 = \frac{A^2}{a}$$

c) $\dot{a} - 1 = \frac{A^2}{a}$ is the Friedmann equation

$$\begin{cases} a(\eta) = \frac{1}{2} A^2 (\cosh \eta - 1) \\ t(\eta) = \frac{1}{2} A^2 (\sinh \eta - \eta) \end{cases}$$

are the parametric equations for the negative curvature models $K = -1$, $\Omega < 1$

η is the conformal time: $dt = a(t) d\eta$

$$\dot{a} = \frac{da}{dt} = \frac{da/d\eta}{dt/d\eta} = \frac{\frac{1}{2} A^2 \sinh \eta}{\frac{1}{2} A^2 (\cosh \eta - 1)} = \frac{\sinh \eta}{\cosh \eta - 1}$$

$$\textcircled{1} \quad \dot{a}^2 - 1 = \frac{\sinh^2 \eta}{(\cosh \eta - 1)^2} - 1 = \frac{\sinh^2 \eta - (\cosh \eta - 1)^2}{(\cosh \eta - 1)^2} =$$

$$= \frac{\sinh^2 \eta - \cosh^2 \eta + 2 \cosh \eta - 1}{(\cosh \eta - 1)^2} \quad (\cosh^2 \eta - \sinh^2 \eta = 1)$$

$$= \frac{-1 + 2 \cosh \eta - 1}{(\cosh \eta - 1)^2} = \frac{2(\cosh \eta - 1)}{(\cosh \eta - 1)^2} = \frac{2}{\cosh \eta - 1}$$

$$\textcircled{2} \quad \frac{A^2}{a} = \frac{A^2}{dt/d\eta} = \frac{A^2}{\frac{1}{2} A^2 (\cosh \eta - 1)} = \frac{2}{\cosh \eta - 1}$$

$\textcircled{1} = \textcircled{2} \rightarrow$ the Friedmann equation is satisfied

d) this model expands from a Big Bang singularity at $\eta = 0$ where $t = 0$, $a = 0$, $\rho = \infty$

the scale factor starts from 0 at $t = 0$ ($\eta = 0$) and increases with no bounds: it means that there is not a big crunch, the universe expands forever.

This is what we expect for an open universe with negative curvature $K = -1$



PROBLEM 3

the variational principle states that the world line of a free particle extremizes the proper time:

a) $\Delta \tau = \int d\tau = \int d\sigma \frac{d\tau}{d\sigma}$ where $L = \frac{d\tau}{d\sigma}$ is the lagrangian

$$L = \frac{d\tau}{d\sigma} = \frac{ds}{d\sigma} = \left[g_{\alpha\beta}(x) \frac{dx^\alpha}{d\sigma} \frac{dx^\beta}{d\sigma} \right]^{\frac{1}{2}}$$

$$ds^2 = \eta^{-2} (dx^2 + dy^2) = g_{\alpha\beta} dx^\alpha dx^\beta \quad (\eta > 0) \text{ is the metric}$$

$$L = \left[\eta^{-2} \left(\frac{dx}{d\sigma} \right)^2 + \eta^{-2} \left(\frac{dy}{d\sigma} \right)^2 \right]^{\frac{1}{2}} = \left\{ \left[\eta^{-2} \left(\left(\frac{dx}{d\sigma} \right)^2 + \left(\frac{dy}{d\sigma} \right)^2 \right) \right] \right\}^{\frac{1}{2}}$$

b) euler - lagrange equations:

$$\textcircled{1} - \frac{d}{d\sigma} \left(\frac{\partial L}{\partial \left(\frac{\partial x}{\partial \sigma} \right)} \right) + \frac{\partial L}{\partial x} = 0$$

$$\frac{\partial L}{\partial \left(\frac{\partial x}{\partial \sigma} \right)} = \frac{1}{2} (L^2)^{-\frac{1}{2}} 2 \eta^{-2} \frac{dx}{d\sigma} \quad \left. \begin{array}{l} \frac{d}{d\sigma} \left[\frac{1}{L \eta^2} \frac{dx}{d\sigma} \right] = 0 \\ \frac{\partial L}{\partial x} = 0 \end{array} \right\}$$

$$L = \frac{ds}{d\sigma} \rightarrow \frac{d}{d\sigma} \left[\frac{1}{\eta^2} \frac{d\sigma}{ds} \frac{dx}{d\sigma} \right] = 0 \quad \text{and} \quad \frac{d}{d\sigma} \left[\frac{1}{\eta^2} \frac{d\sigma}{ds} \right] = 0$$

$$\text{multiplying by } \frac{d\sigma}{ds}: \quad \frac{d}{ds} \left[\frac{1}{\eta^2} \frac{dx}{ds} \right] = 0$$

$$\frac{1}{\eta^2} \frac{d^2 x}{ds^2} - \frac{2}{\eta^3} \frac{dy}{ds} \frac{dx}{ds} = 0 \rightarrow \ddot{x} - \frac{2}{\eta} \dot{x} \dot{y} = 0$$

$$\textcircled{2} - \frac{d}{d\sigma} \left(\frac{\partial L}{\partial \left(\frac{\partial y}{\partial \sigma} \right)} \right) + \frac{\partial L}{\partial y} = 0$$

$$\frac{\partial L}{\partial \left(\frac{\partial y}{\partial \sigma} \right)} = \frac{1}{2} (L^2)^{-\frac{1}{2}} \eta^{-2} 2 \frac{dy}{d\sigma} = \frac{1}{L} \eta^{-2} \frac{dy}{d\sigma}$$

$$\frac{\partial L}{\partial y} = \frac{1}{2} (L^2)^{-\frac{1}{2}} (-2) \eta^{-3} \left[\left(\frac{dx}{d\sigma} \right)^2 + \left(\frac{dy}{d\sigma} \right)^2 \right] = -\frac{1}{L} \eta^{-3} \left[\left(\frac{dx}{d\sigma} \right)^2 + \left(\frac{dy}{d\sigma} \right)^2 \right]$$

$$\frac{d}{d\sigma} \left(\frac{1}{Ly^2} \frac{dy}{d\sigma} \right) + \frac{1}{Ly^3} \left[\left(\frac{dx}{d\sigma} \right)^2 + \left(\frac{dy}{d\sigma} \right)^2 \right] = 0$$

$$L = \frac{ds}{d\sigma}$$

multiplying by $\frac{d\sigma}{ds}$:

$$\frac{d}{ds} \left(\frac{1}{y^2} \frac{dy}{ds} \right) + \frac{1}{y^3} \left(\frac{d\sigma}{ds} \right)^2 \left[\left(\frac{dx}{d\sigma} \right)^2 + \left(\frac{dy}{d\sigma} \right)^2 \right] = 0$$

$$\frac{1}{y^2} \frac{d^2 y}{ds^2} + \left(\frac{dy}{ds} \right)^2 \left(-\frac{2}{y^3} \right) + \frac{1}{y^3} \left[\left(\frac{dx}{ds} \right)^2 + \left(\frac{dy}{ds} \right)^2 \right] = 0$$

$$\left[\frac{d^2 y}{ds^2} - \frac{2}{y} \left(\frac{dy}{ds} \right)^2 + \frac{1}{y} \left(\frac{dx}{ds} \right)^2 + \frac{1}{y} \left(\frac{dy}{ds} \right)^2 \right] = 0$$

$$\frac{d^2 y}{ds^2} - \frac{2}{y} \left(\frac{dy}{ds} \right)^2 + \frac{1}{y} \left(\frac{dx}{ds} \right)^2 + \frac{1}{y} \left(\frac{dy}{ds} \right)^2 = 0$$

$$\ddot{y} + \frac{1}{y} \dot{x}^2 - \frac{1}{y} \dot{y}^2 = 0$$

c) the geodesic equations, from point b) are:

$$\frac{d^2 x}{ds^2} - \frac{2}{y} \frac{dx}{ds} \frac{dy}{ds} = 0 \quad \text{and} \quad \frac{d^2 y}{ds^2} + \frac{1}{y} \left(\frac{dx}{ds} \right)^2 - \frac{1}{y} \left(\frac{dy}{ds} \right)^2 = 0$$

comparing with the general form of the geodesic equation:

$$\frac{d^2 x^\alpha}{d\tau^2} = - \Gamma_{\beta\gamma}^\alpha \frac{dx^\beta}{d\tau} \frac{dx^\gamma}{d\tau} \quad ds = d\tau$$

$$(1) \quad \frac{d^2 x}{ds^2} = \frac{2}{y} \frac{dx}{ds} \frac{dy}{ds} \quad \begin{matrix} \alpha = x \\ \beta = x \\ \gamma = y \end{matrix} \rightarrow \Gamma_{xy}^x = \Gamma_{yx}^x = -\frac{1}{y}$$

$$(2) \quad \frac{d^2 y}{ds^2} = -\frac{1}{y} \frac{dx}{ds} \frac{dx}{ds} + \frac{1}{y} \frac{dy}{ds} \frac{dy}{ds}$$

$$\begin{matrix} \alpha = y \\ \beta = x \\ \gamma = x \end{matrix} \rightarrow \Gamma_{xx}^y = \frac{1}{y}$$

$$\begin{matrix} \alpha = y \\ \beta = y \\ \gamma = y \end{matrix} \rightarrow \Gamma_{yy}^y = -\frac{1}{y}$$

d) the metric is invariant under translations $x \rightarrow x + c$ and $\xi^a = (1, 0)$ is a Killing vector

$$\frac{\partial p_{\alpha\beta}}{\partial x} = 0 \quad \text{and} \quad \frac{\partial \mathcal{L}}{\partial x} = 0$$

Euler Lagrange equations for x

$$-\frac{d}{d\sigma} \left(\frac{\partial \mathcal{L}}{\partial \left(\frac{dx}{d\sigma} \right)} \right) + \frac{\partial \mathcal{L}}{\partial x} = 0$$

$$\rightarrow \frac{d}{d\sigma} \left(\frac{\partial \mathcal{L}}{\partial \left(\frac{dx}{d\sigma} \right)} \right) = 0 \quad \text{and} \quad \frac{\partial \mathcal{L}}{\partial \left(\frac{dx}{d\sigma} \right)} = \text{const.}$$

$$\frac{\partial \mathcal{L}}{\partial \left(\frac{dx}{d\sigma} \right)} = \frac{1}{2} (1)^{\frac{1}{2}} \left(g_{1\beta} \frac{dx^\beta}{d\sigma} \right) = \frac{1}{2L} g_{1\beta} \frac{dx^\beta}{d\tau} \frac{d\tau}{d\sigma}$$

$$= \frac{1}{2} g_{1\beta} \frac{dx^\beta}{d\tau} = \text{const}$$

$$\rightarrow g_{\alpha\beta} \xi^a u^\beta = \text{const}$$

$$g_{\alpha\beta} \xi^a u^\beta = g_{xx} \frac{dx}{d\tau} = \frac{1}{y^2} \frac{dx}{d\tau}$$

this is a conserved quantity and expresses the x -component of the linear momentum.

e) the metric is: $ds^2 = \frac{1}{y^2} dx^2 + \frac{1}{y^2} dy^2$

$$\vec{u} \cdot \vec{u} = 1$$

$$\vec{u} \cdot \vec{u} = g_{\alpha\beta} u^\alpha u^\beta = \frac{1}{y^2} (\dot{x}^2 + \dot{y}^2) = 1 \rightarrow \dot{x}^2 + \dot{y}^2 = y^2$$

the first geodesic equation found in point b is:

$$\ddot{x} - \frac{2}{y} \dot{x} \dot{y} = 0 \rightarrow \frac{\ddot{x}}{\dot{x}} = 2 \frac{\dot{y}}{y}$$

integrating:

$$\ln \dot{x} = 2 \ln y + A \rightarrow \ln \dot{x} = \ln y^2 + A$$

$$\text{and } \dot{x} = A y^2$$

$$\left. \begin{aligned} \dot{x}^2 + \dot{y}^2 &= y^2 \\ \dot{x} &= A y^2 \end{aligned} \right\} \rightarrow A^2 y^4 + \dot{y}^2 = y^2$$

$$\dot{y}^2 = y^2 (1 - A^2 y^2)$$

$$\dot{y} = \pm y (1 - A^2 y^2)^{\frac{1}{2}}$$

$$\frac{dx}{dy} = \frac{dx}{ds} \cdot \frac{ds}{dy} = \frac{\dot{x}}{\dot{y}} = \frac{\pm A y^2}{y \sqrt{1 - A^2 y^2}} = \pm \frac{A y}{\sqrt{\frac{1}{A^2} - y^2}}$$

$$dx = \pm \frac{y}{\sqrt{\frac{1}{A^2} - y^2}} dy$$

$$\int \frac{y dy}{\sqrt{\frac{1}{A^2} - y^2}} = -\sqrt{\frac{1}{A^2} - y^2}$$

$$\text{here } r^2 = \frac{1}{A^2}$$

integrating

$$x = x_0 \mp \sqrt{\frac{1}{A^2} - y^2}$$

$$(x - x_0)^2 = \frac{1}{A^2} - y^2$$

and

$$(x - x_0)^2 + y^2 = \frac{1}{A^2}$$

Since $y > 0$ this equation represents semi-circles in the upper half plane, centered in $(x_0, 0)$ and radius $r = \frac{1}{A}$